The aim of the paper is to give a rigorous derivation of the hydrostatic approximation by taking the small aspect ratio limit to the Navier-Stokes equations. The aspect ratio (the ratio of the depth to horizontal width) is a geometrical constraint in the general large scale geophysical motions meaning that the vertical scale is significantly smaller than horizontal.

Keywords: anisotropic Naiver-Stokes equations, aspect ratio limit, hydrostatic approximation, compressible Primitive Equations.

1 Introduction

The study of behavior of the atmosphere and ocean have attracted attention in the scientific research community. To model the motion and state of the atmosphere the equations of compressible motion is used. In such model the vertical scale of the atmosphere is much more smaller than the planetary horizontal scale. Therefore, many scientists have suggested the viscosity coefficients must be anisotropic, such as [4, 19, 21]. The anisotropic Navier-Stokes equations are widely used in geophysical fluid dynamics. In this paper, we consider the following compressible anisotropic Navier-Stokes equations

\[
\begin{align*}
\rho_t + \text{div}(\rho \mathbf{u}) &= 0, \\
(\rho \mathbf{u})_t + \text{div}(\rho \mathbf{u} \otimes \mathbf{u}) + \nabla p(\rho) &= \mu_x \Delta_x \mathbf{u} + \mu_z \partial_{zz} \mathbf{u},
\end{align*}
\]

in the thin domain \((0, T) \times \Omega \). Here \(\Omega = \{(x, z)| x \in \mathbb{T}^2, -\epsilon < z < \epsilon\}\), \(x\) denotes the horizontal direction and \(z\) denotes the vertical direction, while, \(\mu_x\) and \(\mu_z\) are given constant horizontal viscous coefficient and vertical viscous coefficient. The velocity \(\mathbf{u} = (v, w)\) where \(v(t, x, z) \in \mathbb{R}^2\) and \(w(t, x, z) \in \mathbb{R}\) represent the horizontal velocity and vertical velocity, respectively. Throughout this paper, we use \(\text{div} \mathbf{u} = \text{div}_x \mathbf{v} + \partial_z w\) and \(\nabla = (\nabla_x, \partial_z)\) to denote the three-dimensional spatial divergence and gradient respectively, and \(\Delta_x\) stands for horizontal Laplacian. As atmosphere and ocean are the thin layers, where the fluid layer depth is small compared to radius of sphere, Pedlosky [19] pointed out that “the pressure difference between any two points on the same vertical line depends only on the weight of the fluid between these points...”. Here we neglect the gravity and suppose the pressure \(p(\rho)\) satisfies the barotropic pressure law where the pressure and the density are related by the formula: \(\rho(\rho) = \rho^\gamma\) \((\gamma > 1)\). **Therefore we assume the density \(\rho\) is independent of \(z\) that is \(\rho = \rho(t, x)\).** This plausible assumption agrees well with experiment and is frequently taken as a hypothesis in geophysical fluid dynamics.

Similar to the assumptions by [1, 14], we suppose \(\mu_x = 1\) and \(\mu_z = \epsilon^2\). As stressed by Azérad and Guillén [1], it is necessary to consider the above anisotropic viscosities scaling, which is fundamental for the derivation of Primitive Equations (PE).\(^1\)

\[\begin{align*}
\partial_t \mathbf{v} + \nabla_x (\mathbf{v} \otimes \mathbf{v}) + \partial_z (w \mathbf{v}) + \nabla_x p = \Delta_x \mathbf{v} + \partial_{zz} \mathbf{v}, \\
\partial_t p = 0 \\
\text{div}_x \mathbf{v} + \partial_z w = 0,
\end{align*}\]

\(^1\)For completeness we will write the system (PE):
Under this assumption, the system is rewritten as the following

\begin{align}
\begin{cases}
\partial_t \rho + \text{div}_x (\rho \textbf{v}) + \partial_x (\rho w) = 0,
\rho \partial_t \textbf{v} + \rho (\textbf{u} \cdot \nabla) \textbf{v} - \Delta_x \textbf{v} - \epsilon^2 \partial_{zz} \textbf{v} + \nabla_x p(\rho) = 0,
\rho \partial_t w + \rho u \cdot \nabla w - \Delta_x w - \epsilon^2 \partial_{zz} w + \partial_z p(\rho) = 0.
\end{cases}
\end{align}

(1.3)

Inspired by \cite{1,14}, we introduce the following new unknowns,

\[ u_\epsilon = (v_\epsilon, w_\epsilon), \textbf{v}_\epsilon(x, z, t) = \textbf{v}(x, \epsilon z, t), w_\epsilon = \frac{1}{\epsilon} w(x, \epsilon z, t), \rho_\epsilon = \rho(x, t), \]

for any \((x, z) \in \Omega := \mathbb{T}^2 \times (-1, 1)\). Then the system (1.3) becomes the following compressible scaled Navier-Stokes equations (CNS):

\begin{align}
\begin{cases}
\partial_t \rho_\epsilon + \text{div}_x (\rho_\epsilon \textbf{v}_\epsilon) + \partial_x (\rho_\epsilon w_\epsilon) = 0,
\rho_\epsilon \partial_t \textbf{v}_\epsilon + \rho_\epsilon (\textbf{u}_\epsilon \cdot \nabla) \textbf{v}_\epsilon - \Delta_x \textbf{v}_\epsilon - \partial_{zz} \textbf{v}_\epsilon + \nabla_x p(\rho_\epsilon) = 0,
\epsilon^2 (\rho_\epsilon \partial_t w_\epsilon + \rho_\epsilon \textbf{u}_\epsilon \cdot \nabla w_\epsilon - \Delta_x \textbf{w}_\epsilon - \partial_{zz} w_\epsilon) + \partial_z p(\rho_\epsilon) = 0.
\end{cases}
\end{align}

(1.4)

We supplement the CNS with the following boundary and initial conditions \cite{2}:

\[ \rho_\epsilon, \textbf{u}_\epsilon \text{ are periodic in } x, y, z, \]
\[ (\rho_\epsilon, \textbf{u}_\epsilon)|_{t=0} = (\rho_0, \textbf{u}_0). \]

(1.5)

The goal of this work is to investigate the limit process \(\epsilon \to 0\) in the system of (1.4). Precisely, to show that the system of (1.4) converges in a certain sense to the following compressible Primitive Equations (CPE):

\begin{align}
\begin{cases}
\partial_t \rho + \text{div}_x (\rho \textbf{v}) + \partial_x (\rho w) = 0,
\partial_t (\rho \textbf{v}) + \text{div}_x (\rho \textbf{v} \otimes \textbf{v}) + \partial_x (\rho \textbf{v} \otimes \textbf{v}) + \nabla_x p(\rho) = \Delta_x \textbf{v} + \partial_{zz} \textbf{v},\n\partial_z p(\rho) = 0.
\end{cases}
\end{align}

(1.6)

Our goal is to rigorously show the limit in the framework of weak solutions of CNS. Recently, Bella, Feireisl and Novotný \cite{2}, Malteze and Novotný \cite{18} proved the limit passage from 3D compressible Navier-Stokes equations to 1D and 2D compressible Navier-Stokes equations in thin domain. See also result by Ducomet et al. \cite{6}. Heuristically, inspired by their works, we develop and adapt the corresponding idea of relative entropy inequality for compressible Navier-Stokes equations. **There are significant differences of the mathematical structure between Navier-Stokes equations and CPE model.** Due to the hydrostatic approximation, **there is no information for the vertical velocity in the momentum equation of CPE model, and the vertical velocity is determined by the horizontal velocity via the continuity equation**, so it is very difficult to analyze the CPE model. Therefore, the classical method used in Navier-Stokes system can not be applied straightforwardly to CPE. Luckily, based on our previous work \cite{11} of weak-strong uniqueness to CPE, we prove the aspect ratio limit of compressible anisotropic Navier-Stokes equations. This is the first work to use the relative entropy inequality for proving the hydrostatic approximation at the compressible case. For the introduction of the versatile relative entropy inequality, see \cite{10}. Last but not least, let us mention that the corner-stone analysis of our results is based on the relative energy inequality which was invented by Dafermos, see \cite{5} and by Germain,\cite{13}, who introduced it into compressible Navier-Stokes equations. After that Feireisl and his co-authors \cite{7,8,9} generalized the relative energy inequality and applied such inequality for solving various compressible fluid model problems.

2 Main result

Before stating our main result, we give the definition of a weak solution for CNS and a strong solution for CPE. Recently, Bresch and Jabin \cite{3} introduced different compactness method from

\footnote{Let us mention that explanation of the periodicity boundary conditions for compressible PE can be found in paper by Liu and Titi–arxiv:1905.09367, Page 4, line 14.}
Lions or Feireisl which can be applied to anisotropic stress tensor.\(^3\) Bresch and Jabin has proved the
global existence of weak solutions for non-monotone pressure and for anisotropic stress tensor.
Therefore, following their theory, we can obtain the existence of weak solution for monotone
pressure and anisotropic stress tensor. Let us recall their definitions here.

### 2.1 Dissipative weak solutions of CNS

**Definition 2.1** We say that \([\rho_\epsilon, u_\epsilon]\) with \(u_\epsilon = (v_\epsilon, w_\epsilon)\) is a finite energy weak solution to the
system of \((1.4)\), supplemented with initial data \((1.5)\) if \(\rho_\epsilon = \rho_\epsilon(x, t)\) and
\[
\begin{align*}
(1.4) & \quad \rho_\epsilon \in L^2(0, T; H^1(\Omega)), \quad \rho u_\epsilon^2 \in L^\infty(0, T; L^1(\Omega)), \\
(2.1) & \quad \rho_\epsilon \in L^\infty(0, T; L^2(\Omega)) \cap C([0, T], L^1(\Omega)), \quad \rho_\epsilon \geq 0, \quad \gamma > 4,
\end{align*}
\]

- the continuity equation
\[
\int_\Omega \rho_\epsilon \psi dxdz|_{t=0}^{t=T} = \int_0^T \int_{\Omega} \rho_\epsilon \partial_t \psi + \rho_\epsilon v_\epsilon \cdot \nabla \psi + \rho_\epsilon w_\epsilon \partial_z \psi dxdzdt,
\]
holds for all \(\psi \in C^\infty_c([0, T] \times \Omega);
- the momentum equation
\[
\int_\Omega \rho_\epsilon v_\epsilon \phi_H dxdz|_{t=0}^{t=T} = \int_0^T \int_{\Omega} \rho_\epsilon \partial_t \phi_H + \rho_\epsilon v_\epsilon \partial_x \phi_H + \rho_\epsilon w_\epsilon \partial_z \phi_H + \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \cdot \nabla \phi_H dxdzdt,
\]
and
\[
\begin{align*}
\rho_\epsilon w_\epsilon \partial_3 \phi_3 dxdz|_{t=0}^{t=T} &= \int_0^T \int_{\Omega} \rho_\epsilon w_\epsilon \partial_t \phi_3 + \rho_\epsilon v_\epsilon \partial_x \phi_3 + \rho_\epsilon w_\epsilon \partial_z \phi_3 dxdzdt, \\
\int_\Omega \rho_\epsilon v_\epsilon \phi_{H3} dxdz|_{t=0}^{t=T} &= \int_0^T \int_{\Omega} \rho_\epsilon \partial_t \phi_{H3} + \rho_\epsilon v_\epsilon \partial_x \phi_{H3} + \rho_\epsilon w_\epsilon \partial_z \phi_{H3} + \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \cdot \nabla \phi_{H3} dxdzdt.
\end{align*}
\]

holds for any spatially periodic function \(\phi_H, \phi_3 \in C^\infty_c([0, T] \times \Omega),\) Combining \((2.3)\) and \((2.4)\), we obtain
\[
\begin{align*}
\int_\Omega \rho_\epsilon v_\epsilon \phi_{H3} dxdz + \varepsilon^2 \int_\Omega \rho_\epsilon w_\epsilon \phi_{H3} dxdz|_{t=0}^{t=T} & - \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \partial_t \phi_{H3} dxdzdt - \varepsilon^2 \int_0^T \int_{\Omega} \rho_\epsilon w_\epsilon \partial_t \phi_{H3} dxdzdt \\
& - \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \nabla \phi_{H3} dxdzdt + \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \partial_3 \phi_{H3} dxdzdt - \varepsilon^2 \int_0^T \int_{\Omega} \rho_\epsilon w_\epsilon \partial_3 \phi_{H3} dxdzdt \\
& - \varepsilon^2 \int_0^T \int_{\Omega} \rho_\epsilon v_\epsilon \partial_3 \phi_{H3} dxdzdt + \int_0^T \int_{\Omega} \nabla v_\epsilon \cdot \nabla \phi_{H3} dxdzdt + \int_0^T \int_{\Omega} \nabla w_\epsilon \cdot \nabla \phi_{H3} dxdzdt - \int_0^T \int_{\Omega} \rho_\epsilon |\nabla \phi_{H3}|^2 dxdzdt = 0,
\end{align*}
\]
where spatially periodic function \(\varphi = (\phi_H, \phi_3) \in C^\infty_c([0, T] \times \Omega)\) and \(\text{div} \varphi = \text{div}_x \phi_H + \partial_z \phi_3,\)
- the energy inequality
\[
\begin{align*}
\int_\Omega \frac{1}{2} \rho_\epsilon |v_\epsilon|^2 + \frac{\gamma^2}{2} \rho_\epsilon |w_\epsilon|^2 + P(\rho) \epsilon dxdz|_{t=0}^{t=T} + \int_0^T \int_{\Omega} (|\nabla v_\epsilon|^2 + \varepsilon^2 |\nabla w_\epsilon|^2) dxdzdt & \leq 0,
\end{align*}
\]
holds for a.a \(\tau \in (0, T),\) where \(P(\rho) = \rho \int_{\Omega} \frac{\rho_\epsilon}{\rho} dz.\)

\(^3\)Let us emphasize the result of Bresch and Jabin is valid only for small coefficients of viscosities.
2.2 Strong solution of CPE

We say that \((r, U)\), \(U = (V,W)\) is a strong solution to the CPE system (1.6) in \((0, T) \times \Omega\), if
\[
\begin{align*}
r^2 &\in L^\infty(0, T; H^2(\Omega)), \quad \partial_t r^2 \in L^\infty(0, T; H^1(\Omega)), \quad r > 0 \text{ for all } (t, x), \\
V &\in L^\infty(0, T; H^3(\Omega)) \cap L^2(0, T; H^4(\Omega)), \quad \partial_t V \in L^2(0, T; H^2(\Omega)),
\end{align*}
\]
with initial data \(r_0^2 \in H^2(\Omega), r_0 > 0\) and \(V_0 \in H^3(\Omega)\). Liu and Titi [15] have proved the local existence of strong solution to CPE system (1.6).

**Remark 2.1** As the density is independent of \(z\), we can obtain the following information of vertical velocity for the weak solution of CNS :
\[
\rho w(x, z, t) = -\text{div}_x(\rho \vec{V}) + z \text{div}_x(\rho \nabla), \quad \text{in the sense of } H^{-1}(\Omega),
\]
where
\[
\vec{V}(x, z, t) = \int_0^z v(x, s, t) ds, \quad \Omega(x, t) = \int_0^1 v(x, z, t) dz.
\]

Similarly, we can obtain the same equation for the strong solution of CPE in the classical sense. There is no information about \(w\), so we need to derive its information. We should emphasize that (2.7) is the key step to obtain the existence of weak solution for CPE in [16, 20], which is inspired by incompressible case.

2.3 Versatile relative entropy inequality

Motivated by [7, 8], for any finite energy weak solution \((\rho, u)\), where \(u = (v, w)\), to the CNS system, we introduce the relative energy functional
\[
\mathcal{E}(\rho, u| r, U) = \int_\Omega \left[ \frac{1}{2} |V - V|^2 + \frac{\epsilon^2}{2} |W|^2 + P(\rho) - P'(\rho)(\rho - r) - P(r) \right] dx dz
\]
\[
= \int_\Omega \left[ \frac{1}{2} |V|^2 + \frac{\epsilon^2}{2} |W|^2 + P(\rho) \right] dx dz - \int_\Omega (\rho v \cdot \nabla v + \epsilon^2 \rho w) dx dz
\]
\[
+ \int_\Omega \left[ \frac{|V|^2}{2} + \frac{\epsilon^2}{2} |W|^2 - \rho P'(\rho) \right] dx dz + \int_\Omega p(r) dx dz
\]
\[
= \sum_{i=1}^{4} I_i,
\]
(2.8)

where \(r > 0\), \(U = (V, W)\) are smooth “test” functions, \(V, W \in C^\infty_c([0, T] \times \Omega)\) are spatially periodic functions. Here we have used \(rP'(r) - P(r) = p(r)\).

**Lemma 2.1** Let \((\rho, v, w)\) be a dissipative weak solution introduced in Definition 2.1. Then \((\rho, v, w)\) satisfy the versatile relative entropy inequality
\[
\mathcal{E}(\rho, u| r, U)|_{t=0}^{\tau} + \int_0^\tau \int_\Omega (\nabla v : (\nabla v - \nabla V) + \epsilon^2 |\nabla w|^2) dx dz dt
\]
\[
\leq \int_0^\tau \int_\Omega \rho (\partial_t V + v \cdot \nabla_x V + w \partial_x V)(V - v) dx dt
\]
\[
+ \epsilon^2 \int_0^\tau \int_\Omega \rho (\partial_t W + v \cdot \nabla_x W + w \partial_x W)(W - w) dx dz dt + \epsilon^2 \int_0^\tau \int_\Omega \nabla v \cdot \nabla W dx dz dt
\]
\[
- \int_0^\tau \int_\Omega \rho''(r)((\rho - r) \partial_t r + \rho v \cdot \nabla_x r) dx dz dt - \int_0^\tau \int_\Omega p(r) dx dz dt.
\]
(2.9)

\footnote{The calculation of \(\rho w\) can be seen in details see Liu and Titi [17] page 1920 and see [20], page 4.}
2.4 Main result

Now, we are ready to state our main result.

**Theorem 2.1** Let $\gamma > 4$, $T_{\text{max}} > 0$ be the life time of strong solution to CPE system (1.6) corresponding to initial data $[r_0, V_0]$. Let $(\rho_\epsilon, u_\epsilon) = (v_\epsilon, w_\epsilon)$ be a sequence of dissipative weak solutions to the CNS system (1.4) from the initial data $(\rho_{0,\epsilon}, u_{0,\epsilon})$. Suppose that $\rho_{0,\epsilon} > 0$, $r_0 > 0$ and

$$E(\rho_{0,\epsilon}, u_{0,\epsilon} | r_0, U_0) \to 0,$$

where $U_0 = (V_0, W_0)$, then

$$\text{ess sup}_{t \in (0, T_{\text{max}})} E(\rho_\epsilon, u_\epsilon | r, U) \to 0,$$

where $U = (V, W)$ and the couple $(r, U)$ satisfy the CPE system (1.6) on the time interval $[0, T_{\text{max}}]$.

For more details, see [10, 12].

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